Physics Letters B 791 (2019) 236-241

Contents lists available at ScienceDirect

Physics Letters B

www.elsevier.com/locate/physletb

DREENA-B framework: First predictions of R_{AA} and v_2 within dynamical energy loss formalism in evolving QCD medium

Dusan Zigic^a, Igor Salom^a, Jussi Auvinen^a, Marko Djordjevic^b, Magdalena Djordjevic^{a,*}

^a Institute of Physics Belgrade, University of Belgrade, Serbia

^b Faculty of Biology, University of Belgrade, Serbia

ARTICLE INFO

Article history: Received 16 July 2018 Received in revised form 10 December 2018 Accepted 15 February 2019 Available online 25 February 2019 Editor: W. Haxton

Keywords: Relativistic heavy ion collisions Quark-gluon plasma LHC Heavy flavor suppression High pt hadrons

ABSTRACT

Dynamical energy loss formalism allows generating state-of-the-art suppression predictions in finite size QCD medium, employing a sophisticated model of high- p_{\perp} parton interactions with QGP. We here report a major step of introducing medium evolution in the formalism though 1 + 1D Bjorken ("B") expansion, while preserving all complex features of the original dynamical energy loss framework. We use this framework to provide joint R_{AA} and v_2 predictions, for the first time within the dynamical energy loss formalism in evolving QCD medium. The predictions are generated for a wide range of high p_{\perp} observables, i.e. for all types of probes (both light and heavy) and for all centrality regions in both Pb + Pb and Xe + Xe collisions at the LHC. Where experimental data are available, DREENA-B framework leads to a good joint agreement with v_2 and R_{AA} data. Such agreement is encouraging, i.e. may lead us closer to resolving v_2 puzzle (difficulty of previous models to jointly explain R_{AA} and v_2 data), though this still remains to be thoroughly tested by including state-of-the-art medium evolution within DREENA framework. While introducing medium evolution significantly changes v_2 predictions, R_{AA} predictions remain robust and moreover in a good agreement with the experimental data; R_{AA} observable is therefore suitable for calibrating parton-medium interaction model, independently from the medium evolution. Finally, for heavy flavor, we observe a strikingly similar signature of the dead-cone effect on both R_{AA} and v_2 - we also provide a simple analytical understanding behind this result. Overall, the results presented here indicate that DREENA framework is a reliable tool for QGP tomography. © 2019 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license

(http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.

1. Introduction

It is by now established that quark-gluon plasma (QGP), being a new state of matter [1,2] consisting of interacting quarks, antiquarks and gluons, is created in ultra-relativistic heavy ion collisions at the Relativistic Heavy Ion Collider (RHIC) and the Large Hadron Collider (LHC). Energy loss of rare high p_{\perp} particles, which are created in such collisions and which transverse QGP, is considered to be an excellent probe of this form of matter [3–6]. Such energy loss is reflected through different observables, most importantly angular averaged (R_{AA}) [7–14] and angular differential (v_2) [15–22] nuclear modification factor, which can be measured and predicted for both light and heavy flavor probes. Therefore, comparing a comprehensive set of predictions, created under the same model and parameter set, with the corresponding experi-

* Corresponding author.

E-mail address: magda@ipb.ac.rs (M. Djordjevic).

mental data, allows for systematical investigation of QCD medium properties, i.e. QGP tomography.

We previously showed that the dynamical energy loss formalism [23-25] provides an excellent tool for such tomography. In particular, we demonstrated that the formalism shows a very good agreement [27–30] with a wide range of R_{AA} data, coming from different experiments, collision energies, probes and centralities. Recently, we also used this formalism to generate first v_2 predictions, within DREENA-C framework [26], where DREENA stands for Dynamical Radiative and Elastic ENergy loss Approach, and "C" denotes constant temperature QCD medium. These predictions were compared jointly with R_{AA} and v_2 data, showing a very good agreement with R_{AA} data, while visibly overestimating v_2 data. This overestimation also clearly differentiates the dynamical energy loss from other models, which systematically underestimated the v_2 data, leading to the so called v_2 puzzle [31–33]. On the other hand, it is also clear that v_2 predictions have to be further improved - in particular v_2 was shown to be sensitive to medium evolution, while in DREENA-C medium evolution

https://doi.org/10.1016/j.physletb.2019.02.020

0370-2693/© 2019 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.







was introduced in the simplest form, through constant medium temperature. This problem then motivated us to introduce medium evolution in DREENA framework.

While several existing energy loss approaches already contain a sophisticated medium evolution, they employ simplified parton energy loss models. On the other hand, our dynamical energy loss formalism corresponds to the opposite "limit", where constant (mean) medium temperature was assumed, combined with a sophisticated model of parton-medium interactions, which includes: *i*) QCD medium composed of dynamical (i.e. moving) scattering centers, which is contrary to the widely used static scattering centers approximation, *ii*) finite size QCD medium, *iii*) finite temperature QCD medium, modeled by generalized HTL approach [34, 35], naturally regularizing all infrared and ultraviolet divergencies [23–25,36]. *iv*) collisional [25] and radiative [23] energy losses, calculated within the same theoretical framework, *v*) finite parton mass, making the formalism applicable to both light and heavy flavor, *vi*) finite magnetic [37] mass and running coupling [27].

Note that we have previously showed that all the ingredients stated above are important for accurately describing experimental data [38]. Consequently, introducing medium evolution in the dynamical energy loss, is a major step in the model development, as all components in the model have to be preserved, and no additional simplifications should be used in the numerical procedure. In addition to developing the energy loss expressions with changing temperature, we also wanted to develop a framework that can efficiently generate a set of predictions for all types of probes and all centrality regions. That is, we think that for a model to be realistically compared with experimental data, the comparison should be done for a comprehensive set of light and heavy flavor experimental data, through the same numerical framework and the same parameter set. To implement this principle, we also had to develop a numerical framework that can efficiently (i.e. in a reasonably short time frame) generate such predictions, which will be presented in this paper.

We will start the task of introducing the medium evolution in the dynamical energy loss formalism with DREENA-B framework presented here, where "B" stands for Bjorken. In this framework, OCD medium is modeled by the ideal hydrodynamical 1 + 1DBiorken expansion [39], which has a simple analytical form of temperature (T) dependence. This simple T dependence will be used as an intermediate between constant (mean) temperature DREENA-C framework and the full evolution QGP tomography tool. While, on one hand, inclusion of Bjorken expansion in DREENA framework is a major task (having in mind complexity of our model, see above), it on the other hand significantly simplifies the numerical procedure compared to full medium evolutions. This will then allow step-by-step development of full QGP tomography framework, and assessing improvements in the predictions when, within the same theoretical framework, one is transitioning towards more complex QGP evolution models within the dynamical energy loss framework.

2. Computational framework

To calculate the quenched spectra of hadrons, we use the generic pQCD convolution, while the assumptions are provided in [27]:

$$\frac{E_f d^3 \sigma}{dp_f^3} = \frac{E_i d^3 \sigma(Q)}{dp_i^3} \otimes P(E_i \to E_f) \\ \otimes D(Q \to H_Q) \otimes f(H_Q \to e, J/\psi), \tag{1}$$

where "i" and "f", respectively, correspond to "initial" and "final", Q denotes quarks and gluons (partons). $E_i d^3 \sigma(Q)/dp_i^3$ denotes

the initial parton spectrum, computed at next to leading order [40] for light and heavy partons. $D(Q \rightarrow H_Q)$ is the fragmentation function of parton Q to hadron H_Q ; for charged hadrons, D and B mesons we use DSS [41], BCFY [42] and KLP [43] fragmentation functions, respectively. $P(E_i \rightarrow E_f)$ is the energy loss probability, generalized to include both radiative and collisional energy loss in a realistic finite size dynamical QCD medium in which the temperature is changing, as well as running coupling, path-length and multi-gluon fluctuations. In below expressions, running coupling is introduced according to [27], where the temperature T now changes with proper time τ ; the temperature dependence along the jet path is taken according to the ideal hydrodynamical 1 + 1DBjorken expansion [39]. Partons travel different paths in the QCD medium, which is taken into account through path length fluctuations [44]. Multi-gluon fluctuations take into account that the energy loss is a distribution, and are included according to [27,45] (for radiative energy loss) and [44,46] (for collisional energy loss).

The dynamical energy loss formalism was originally developed for constant temperature QCD medium, as described in detail in [23–25]. We have now derived collisional and radiative energy loss expressions for the medium in which the temperature is changing along the jet path; detailed calculations will be presented elsewhere, while the main results are summarized below.

For the collisional energy loss, we obtain the following analytical expression:

$$\frac{dE_{col}}{d\tau} = \frac{2C_R}{\pi v^2} \alpha_S(ET) \alpha_S(\mu_E^2(T))
\int_0^\infty n_{eq}(|\vec{\mathbf{k}}|, T) d|\vec{\mathbf{k}}| \left(\int_0^{|\vec{\mathbf{k}}|/(1+v)} d|\vec{\mathbf{q}}| \int_{-v|\vec{\mathbf{q}}|}^{v|\vec{\mathbf{q}}|} \omega d\omega \right)
+ \int_{|\vec{\mathbf{k}}|/(1+v)}^{|\vec{\mathbf{q}}|-2|\vec{\mathbf{k}}|} d|\vec{\mathbf{q}}| \int_{-v|\vec{\mathbf{q}}|^2}^{v|\vec{\mathbf{q}}|} \omega d\omega \right)
\left(|\Delta_L(q, T)|^2 \frac{(2|\vec{\mathbf{k}}| + \omega)^2 - |\vec{\mathbf{q}}|^2}{2}
+ |\Delta_T(q, T)|^2 \frac{(|\vec{\mathbf{q}}|^2 - \omega^2)((2|\vec{\mathbf{k}}| + \omega)^2 + |\vec{\mathbf{q}}|^2)}{4|\vec{\mathbf{q}}|^4} (v^2|\vec{\mathbf{q}}|^2 - \omega^2) \right).$$
(2)

Here *E* is initial jet energy, τ is the proper time, *T* is the temperature of the medium, α_S is running coupling [27] and $C_R = \frac{4}{3}$. *k* is the 4-momentum of the incoming medium parton, *v* is the velocity of the incoming jet and $q = (\omega, \vec{\mathbf{q}})$ is the 4-momentum of the gluon. $n_{eq}(|\vec{\mathbf{k}}|, T) = \frac{N}{e^{|\vec{\mathbf{k}}|/T} + 1} + \frac{N_f}{e^{|\vec{\mathbf{k}}|/T} + 1}$ is the equilibrium momentum distribution [47] at temperature *T* including quarks and gluons (*N* and N_f are the number of colors and flavors, respectively). $\Delta_L(T)$ and $\Delta_T(T)$ are effective longitudinal and transverse gluon propagators [48]:

$$\Delta_L^{-1}(T) = \vec{\mathbf{q}}^2 + \mu_E(T)^2 (1 + \frac{\omega}{2|\vec{\mathbf{q}}|} \ln |\frac{\omega - |\mathbf{q}|}{\omega + |\vec{\mathbf{q}}|}|),$$
(3)
$$\Delta_T^{-1}(T) = \omega^2 - \vec{\mathbf{q}}^2 - \frac{\mu_E(T)^2}{2}$$

. → .

$$-\frac{(\omega^2 - \vec{\mathbf{q}}^2)\mu_E(T)^2}{2\vec{\mathbf{q}}^2}(1 + \frac{\omega}{2|\vec{\mathbf{q}}|}\ln|\frac{\omega - |\vec{\mathbf{q}}|}{\omega + |\vec{\mathbf{q}}|}|), \qquad (4)$$

while the electric screening (the Debye mass) $\mu_E(T)$ can be obtained by self-consistently solving the expression [49] (n_f is num-

ber of the effective degrees of freedom, Λ_{QCD} is perturbative QCD scale):

$$\frac{\mu_E(T)^2}{\Lambda_{QCD}^2} \ln\left(\frac{\mu_E(T)^2}{\Lambda_{QCD}^2}\right) = \frac{1 + n_f/6}{11 - 2/3 n_f} \left(\frac{4\pi T}{\Lambda_{QCD}}\right)^2.$$
 (5)

The gluon radiation spectrum takes the following form:

$$\frac{dN_{\rm rad}}{dxd\tau} = \int \frac{d^2k}{\pi} \frac{d^2q}{\pi} \frac{2C_R C_2(G) T}{x} \frac{\alpha_s(ET) \alpha_s(\frac{k^2 + \chi(T)}{x})}{\pi} \\
\times \frac{\mu_E(T)^2 - \mu_M(T)^2}{(q^2 + \mu_M(T)^2)(q^2 + \mu_E(T)^2)} \\
\times \left(1 - \cos\frac{(k+q)^2 + \chi(T)}{xE^+} \tau\right) \frac{(k+q)}{(k+q)^2 + \chi(T)} \\
\times \left(\frac{(k+q)}{(k+q)^2 + \chi(T)} - \frac{k}{k^2 + \chi(T)}\right),$$
(6)

where $C_2(G) = 3$ and $\mu_M(T)$ is magnetic screening. k and q are transverse momenta of radiated and exchanged (virtual) gluon, respectively. $\chi(T) \equiv M^2 x^2 + m_E(T)^2/2$, where x is the longitudinal momentum fraction of the jet carried away by the emitted gluon, M is the mass of the quark or gluon jet and $m_g(T) = \mu_E(T)/\sqrt{2}$ is effective gluon mass in finite temperature QCD medium [36]. We also recently abolished the soft-gluon approximation [50], for which we however showed that it does not significantly affect the model results; consequently, this improvement is not included in DREENA-B, but can be straightforwardly implemented in the future DREENA developments, if needed.

Note that, as a result of introducing medium evolution, the dynamical energy loss formalism now explicitly contains changing temperature in the energy loss expression. This is contrary to most of the other models, in which temperature evolution is introduced indirectly, through transport coefficient \hat{q} or gluon rapidity density $\frac{dNg}{dy}$ (see [51] and references therein). This feature makes the dynamical energy loss a natural framework to incorporate diverse temperature profiles as a starting point for QGP tomography. As the first (major) step, we will below numerically implement this possibility through Bjorken 1 + 1*D* expansion [39].

Regarding the numerical procedure, computation efficiency of the algorithm implemented in DREENA-C framework [26] was already two orders of magnitude higher with respect to the basic (unoptimized) brute-force approach applied in [27]. However, straightforward adaptation of the DREENA-C code to the case of the Bjorken type evolving medium was not sufficient. This was dominantly due to additional integration over proper time τ , which increased the calculation time for more than two orders of magnitude. The computation of e.g. radiative energy losses alone, for a single probe, took around 10 hours on the available computer resources (a high performance workstation). Taking into account that it requires ~ 10^3 such runs to produce the results presented in this paper, it is evident that a substantial computational speedup was necessary.

The main algorithmic tool that we used to optimize the calculation was a combination of sampling and tabulating various intermediate computation values and their subsequent interpolation. We used nonuniform adaptive grids of the sampling points, denser in the parts of the parameter volume where the sampled function changed rapidly. Similarly, the parameters used for the numerical integration (the number of quasi-Monte Carlo sampling points and the required accuracy) were also suitably varied throughout the parameter space. Finally, while the computation in DREENA-C was performed in a software for symbolic computation, the new algorithm was redeveloped in C programming language. The combined effect of all these improvements was a computational speedup of almost three orders of magnitude, which was a necessary prerequisite for both current practical applicability and future developments of DREENA framework.

Regarding the parameters, we implement Bjorken 1 + 1D expansion [39], with commonly used $\tau_0 = 0.6$ fm [52,53], and initial temperatures for different centralities calculated according to $T_0 \sim (dN_{ch}/dy/A_{\perp})^{1/3}$ [54], where dN_{ch}/dy is charged multiplicity and A_{\perp} is overlap area (based on the Glauber model nuclear overlap function) for specific collision system and centrality. We use this equation, starting from $T_0 = 500$ MeV in 5.02 TeV Pb + Pbmost central collisions at the LHC, which is estimated based on average medium temperature of 348 MeV in these collisions, and QCD transition temperature of $T_c \approx 150$ MeV [55]. Note that the average medium temperature of 348 MeV in most central 5.02 TeV Pb + Pb collisions comes from [28] the effective temperature (T_{eff}) of 304 MeV for 0-40% centrality 2.76 TeV Pb+Pb collisions at the LHC [56] experiments (as extracted by ALICE). Once T_0 s for most central Pb + Pb collisions are fixed, T_0 for both different centralities and different collision systems (Xe + Xe and Pb + Pb) are obtained from the expression above.

Other parameters used in the calculation remain the same as in DREENA-C [26]. In particular, the path-length distributions for both Xe + Xe and Pb + Pb are calculated following the procedure described in [57], with an additional hard sphere restriction $r < R_A$ in the Woods-Saxon nuclear density distribution to regulate the path lengths in the peripheral collisions. Note that the path-length distributions for Pb + Pb are explicitly provided in [26]; we have also checked that, for each centrality, our obtained eccentricities remain within the standard deviation of the corresponding Glauber Monte Carlo results [58]. For Xe + Xe, it is straightforward to show that Xe + Xe and Pb + Pb distributions are the same up to rescalling factor ($A^{1/3}$, where A is atomic mass number), as we discussed in [59]. Furthermore, the path-length distributions correspond to geometric quantity, and are therefore the same for all types of partons (light and heavy). For QGP, we take $\Lambda_{0,CD} = 0.2$ GeV and $n_f = 3$. As noted above, temperature dependent Debye mass $\mu_E(T)$ is obtained from [49]. For light quarks and gluons, we, respectively, assume that their effective masses are $M \approx \mu_E(T)/\sqrt{6}$ and $m_g \approx \mu_E(T)/\sqrt{2}$ [36]. The charm and bottom masses are M = 1.2GeV and M = 4.75 GeV, respectively. Magnetic to electric mass ratio is extracted from non-perturbative calculations [60,61], leading to $0.4 < \mu_M/\mu_E < 0.6$ - this range of screening masses leads to presented uncertainty in the predictions. We note that no fitting parameters are used in the calculations, that is, all the parameters correspond to standard literature values.

3. Results and discussion

In this section, we will present joint R_{AA} and v_2 predictions for light (charged hadrons) and heavy (D and B mesons) flavor in Pb + Pb and Xe + Xe collisions at the LHC, obtained by DREENA-B framework. Based on the path-length distributions from Figure 1 in [26], we will, in Figs. 1 to 2, show R_{AA} and v_2 predictions for light and heavy flavor, in 5.02 TeV Pb + Pb and 5.44 TeV Xe + Xecollisions, at different centralities. We start by presenting charged hadrons predictions, where R_{AA} data are available for both Pb + Pband Xe + Xe, while v_2 data exist for Pb + Pb collisions. Comparison of our joint predictions with experimental data is shown in Fig. 1, where 1st and 2nd columns correspond, respectively, to R_{AA} and v_2 predictions at Pb + Pb, while 3rd and 4th columns present equivalent predictions/data for Xe + Xe collisions at the LHC. From this figure, we see that DREENA-B is able to well explain joint R_{AA} and v_2 data. For 5.44 TeV Xe + Xe collisions at the LHC, we observe good agreement of our predictions with prelim-



Fig. 1. *First column:* R_{AA} vs. p_{\perp} predictions are compared with 5.02 TeV Pb + Pb ALICE [7], ATLAS [8] and CMS [9] h^{\pm} experimental data. *Second column:* Equivalent comparison for v_2 vs. p_{\perp} (data [15–17]). *Third column:* R_{AA} vs. p_{\perp} predictions are compared with 5.44 TeV Xe + Xe ALICE [62], ATLAS [63] and CMS [64] preliminary data. *Fourth column:* Equivalent predictions for v_2 vs. p_{\perp} . ALICE, ATLAS and CMS data are respectively represented by red circles, green triangles and blue squares, while centrality regions are indicated in the relevant subfigures. Full and dashed curves correspond to, respectively, DREENA-B and DREENA-C frameworks. The gray band boundaries correspond to $\mu_M/\mu_E = 0.4$ and $\mu_M/\mu_E = 0.6$.



Fig. 2. *First column:* Theoretical predictions for D and B meson R_{AA} vs. p_{\perp} are compared with the available 5.02 TeV Pb + Pb ALICE [10] (red circles) D meson experimental data. *Second column:* v_2 vs. p_{\perp} predictions are compared with 5.02 TeV Pb + Pb ALICE [19] (red circles) and CMS [20] (blue squares) D meson experimental data. *Third and fourth column:* Heavy flavor R_{AA} and v_2 vs. p_{\perp} predictions are, respectively, provided for 5.44 TeV Xe + Xe collisions at the LHC. First to third row, respectively, correspond to 0 - 10%, 10 - 30% and 30 - 50% centrality regions. The gray band boundaries correspond to $\mu_M/\mu_E = 0.4$ and $\mu_M/\mu_E = 0.6$.

inary R_{AA} data from ALICE, ATLAS and CMS data (where we note that these predictions were generated, and posted on arXiv, before the data became available), except for high centrality regions, where our predictions do not agree with ALICE (and also partially with ATLAS) data; however, note that in these regions ALICE, ATLAS and CMS data also do not agree with each other.

Furthermore, comparison of predictions obtained with DREENA-B and DREENA-C frameworks in Fig. 1, allows to directly assess the importance of inclusion of medium evolution on different observables, as the main difference between these two frameworks is that DREENA-B contains Bjorken evolution, while DREENA-C accounts for evolution in the simplest form (through constant mean temperature). We see that inclusion of Bjorken evolution has a negligible effect on R_{AA} , while having a significant effect on v_2 . That is, it keeps R_{AA} almost unchanged, while significantly decreasing v_2 . Consequently, small effect on R_{AA} , supports the fact that R_{AA} is weakly sensitive to medium evolution, making R_{AA} an excellent probe of jet-medium interactions in QGP; i.e. in QGP tomography, R_{AA} can be used to calibrate parton medium interaction models. On the other hand, medium evolution clearly influences v_2 predictions, in line with previous conclusions [65,66]; this sensitivity makes v_2 an ideal probe to constrain QGP medium parameters also from the point of high p_{\perp} measurements (in addition to constraining them from low p_{\perp} predictions and data).

In Fig. 2, we provide joint predictions for D and B meson R_{AA} (left panel) and v_2 (right panel) predictions for both 5.02 TeV Pb + Pb and 5.44 TeV Xe + Xe collisions at the LHC. Predictions are compared with the available experimental data. For D mesons, we again observe good joint agreement with the available R_{AA} and v_2 data. For B mesons (where the experimental data are yet to become available), we predict a notable suppression (see also [27,

67]), which is consistent with non-prompt J/Ψ R_{AA} measurements [68] (indirect probe of b quark suppression). Additionally, we predict non-zero v_2 for higher centrality regions. This does not necessarily mean that heavy B meson flows, since we here show predictions for high p_{\perp} , and flow is inherently connected with *low* p_{\perp} v_2 . On the other hand, high p_{\perp} v_2 is connected with the difference in the B meson suppression for different (in-plane and out-of-plane) directions, leading to our predictions of non zero v_2 for *high* p_{\perp} B mesons. Additionally, by comparing D and B meson v_2 s in Fig. 2, we observe that their difference is large and that it qualitatively exhibits the same dependence on p_{\perp} as R_{AA} . This v_2 comparison therefore presents additional important prediction of the heavy flavor dead-cone effect in QGP, where a strikingly similar signature of this effect is observed for R_{AA} and v_2 .

The predicted similarity between R_{AA} and v_2 dead-cone effects can be analytically understood by using simple scaling arguments. Fractional energy loss can be estimated as [26]

$$\Delta E/E \sim \eta T^a L^b,\tag{7}$$

where *a*, *b* are proportionality factors, *T* and *L* are, respectively, the average temperature of the medium and the average path-length traversed by the jet. η is a proportionality factor that depends on initial jet mass *M* and transverse momentum p_{\perp} .

Under the assumption of small fractional energy loss, we can make the following estimate [26]:

$$R_{AA} \approx 1 - \xi(M, p_{\perp})T^{a}L^{b},$$

$$v_{2} \approx \xi(M, p_{\perp}) \frac{(T^{a}L^{b-1}\Delta L - T^{a-1}L^{b}\Delta T)}{2},$$
(8)

where ΔL and ΔT are, respectively, changes in average pathlengths and average temperatures along out-of-plane and in-plane directions. $\xi = (n - 2)\eta/2$, where *n* is the steepness of the initial momentum distribution function.

The difference between R_{AA} and v_2 for D and B mesons then becomes:

$$R^{B}_{AA} - R^{D}_{AA} \approx (\xi(M_{c}, p_{\perp}) - \xi(M_{b}, p_{\perp})) T^{a}L^{b},$$

$$v^{D}_{2} - v^{B}_{2} \approx (\xi(M_{c}, p_{\perp}) - \xi(M_{b}, p_{\perp})))$$

$$\times \frac{(T^{a}L^{b-1}\Delta L - T^{a-1}L^{b}\Delta T)}{2},$$
(9)

where M_c and M_b are charm and bottom quark masses respectively. From Eq. (9), we see the same mass dependent prefactor for both R_{AA} and v_2 comparison, intuitively explaining our predicted dead-cone effect similarity for high- p_{\perp} R_{AA} and v_2 .

4. Summary

Overall, we see that comprehensive joint R_{AA} and v_2 predictions, obtained with our DREENA-B framework, lead to a good agreement with all available light and heavy flavor data. This is, to our knowledge, the first study to provide such comprehensive predictions for high p_{\perp} observables. In the context of v_2 puzzle, this study presents a significant development, as the other models were not able to achieve this agreement without introducing new phenomena [69]. However, for more definite conclusions, the inclusion of more complex QGP evolution within DREENA framework is needed, which is our main ongoing - but highly non-trivial - task, due to the complexity of underlying energy loss formalism.

As an outlook, for Xe + Xe, we also showed an extensive set of predictions for both R_{AA} and v_2 , for different flavors and centralities, to be compared with the upcoming experimental data. Reasonable agreement with these data would present a strong argument that the dynamical energy loss formalism can provide a reliable tool for precision QGP tomography. Moreover, such comparison between predictions and experimental data can also confirm interesting new patterns in suppression data, such as our prediction of strikingly similar signature of the dead-cone effect between R_{AA} and v_2 data.

Acknowledgements

We thank Bojana Blagojevic and Pasi Huovinen for useful discussions. We thank ALICE, ATLAS and CMS Collaborations for providing the shown data. This work is supported by the European Research Council, grant ERC-2016-COG: 725741, and by the Ministry of Science and Technological Development of the Republic of Serbia, under project numbers ON171004, ON173052 and ON171031.

References

- [1] J.C. Collins, M.J. Perry, Phys. Rev. Lett. 34 (1975) 1353.
- [2] G. Baym, S.A. Chin, Phys. Lett. B 62 (1976) 241.
- [3] M. Gyulassy, L. McLerran, Nucl. Phys. A 750 (2005) 30.
- [4] E.V. Shuryak, Nucl. Phys. A 750 (2005) 64;
- E.V. Shuryak, Rev. Mod. Phys. 89 (2017) 035001.
- [5] B. Jacak, P. Steinberg, Phys. Today 63 (2010) 39.
- [6] B. Muller, J. Schukraft, B. Wyslouch, Annu. Rev. Nucl. Part. Sci. 62 (2012) 361.
- [7] S. Acharya, et al., ALICE Collaboration, J. High Energy Phys. 1811 (2018) 013.
- [8] ATLAS Collaboration, ATLAS-CONF-2017-012.
- [9] V. Khachatryan, et al., CMS Collaboration, J. High Energy Phys. 1704 (2017) 039. [10] S. Jaelani, ALICE Collaboration, Int. J. Mod. Phys. Conf. Ser. 46 (2018) 1860018.

- [11] B. Abelev, et al., ALICE Collaboration, Phys. Lett. B 720 (2013) 52.
- [12] B. Abelev, et al., ALICE Collaboration, J. High Energy Phys. 1209 (2012) 112.
- [13] A. Adare, et al., PHENIX Collaboration, Phys. Rev. Lett. 101 (2008) 232301; A. Adare, et al., Phys. Rev. C 87 (3) (2013) 034911.
- [14] B.I. Abelev, et al., STAR Collaboration, Phys. Lett. B 655 (2007) 104.
- [15] S. Acharya, et al., ALICE Collaboration, J. High Energy Phys. 1807 (2018) 103.
- [16] M. Aaboud, et al., ATLAS Collaboration, Eur. Phys. J. C 78 (12) (2018) 997.
- [17] A.M. Sirunyan, et al., CMS Collaboration, Phys. Lett. B 776 (2018) 195.
- [18] B.B. Abelev, et al., ALICE Collaboration, Phys. Rev. C 90 (3) (2014) 034904.
- [19] S. Acharya, et al., ALICE Collaboration, Phys. Rev. Lett. 120 (10) (2018) 102301.
- [20] A.M. Sirunyan, et al., CMS Collaboration, Phys. Rev. Lett. 120 (20) (2018) 202301
- [21] G. Aad, et al., ATLAS Collaboration, Phys. Lett. B 707 (2012) 330.
- [22] S. Chatrchyan, et al., CMS Collaboration, Phys. Rev. Lett. 109 (2012) 022301.
- [23] M. Djordjevic, Phys. Rev. C 80 (2009) 064909.
- [24] M. Djordjevic, U. Heinz, Phys. Rev. Lett. 101 (2008) 022302.
- [25] M. Djordjevic, Phys. Rev. C 74 (2006) 064907.
- [26] D. Zigic, I. Salom, J. Auvinen, M. Djordjevic, M. Djordjevic, arXiv:1805.03494 [nucl-th].
- [27] M. Djordjevic, M. Djordjevic, Phys. Lett. B 734 (2014) 286.
- [28] M. Djordjevic, M. Djordjevic, B. Blagojevic, Phys. Lett. B 737 (2014) 298.
- [29] M. Djordjevic, M. Djordjevic, Phys. Rev. C 92 (2015) 024918.
- [30] M. Djordjevic, Phys. Rev. Lett. 734 (2014) 286;
- M. Djordjevic, Phys. Lett. B 763 (2016) 439.
- [31] J. Noronha-Hostler, B. Betz, J. Noronha, M. Gyulassy, Phys. Rev. Lett. 116 (25) (2016) 252301.
- [32] B. Betz, M. Gyulassy, J. High Energy Phys. 1408 (2014) 090; B. Betz, M. Gyulassy, J. High Energy Phys. 1410 (2014) 043, Erratum.
- [33] D. Molnar, D. Sun, arXiv:1305.1046 [nucl-th].
- [34] J.I. Kapusta, Finite-Temperature Field Theory, Cambridge University Press, 1989.
- [35] M. Le Bellac, Thermal Field Theory, Cambridge University Press, 1996.
- [36] M. Djordjevic, M. Gyulassy, Phys. Rev. C 68 (2003) 034914.
- [37] M. Djordjevic, Phys. Lett. B 709 (2012) 229.
- [38] B. Blagojevic, M. Djordjevic, J. Phys. G 42 (2015) 075105.
- [39] J.D. Bjorken, Phys. Rev. D 27 (1983) 140.
- [40] Z.B. Kang, I. Vitev, H. Xing, Phys. Lett. B 718 (2012) 482;
- R. Sharma, I. Vitev, B.W. Zhang, Phys. Rev. C 80 (2009) 054902. [41] D. de Florian, R. Sassot, M. Stratmann, Phys. Rev. D 75 (2007) 114010.
- [42] M. Cacciari, P. Nason, J. High Energy Phys. 0309 (2003) 006;
- E. Braaten, K.-M. Cheung, S. Fleming, T.C. Yuan, Phys. Rev. D 51 (1995) 4819.
- [43] V.G. Kartvelishvili, A.K. Likhoded, V.A. Petrov, Phys. Lett. B 78 (1978) 615.
- [44] S. Wicks, W. Horowitz, M. Djordjevic, M. Gyulassy, Nucl. Phys. A 784 (2007) 426
- [45] M. Gyulassy, P. Levai, I. Vitev, Phys. Lett. B 538 (2002) 282.
- [46] G.D. Moore, D. Teaney, Phys. Rev. C 71 (2005) 064904.
- [47] E. Braaten, M.H. Thoma, Phys. Rev. D 44 (1991) 1298.
- [48] A.V. Selikhov, M. Gyulassy, Phys. Lett. B 316 (1993) 373;
- A.V. Selikhov, M. Gyulassy, Phys. Rev. C 49 (1994) 1726.
- [49] A. Peshier, arXiv:hep-ph/0601119, 2006.
- [50] B. Blagojevic, M. Djordjevic, M. Djordjevic, Phys. Rev. C 99 (2) (2019) 024901.
- [51] K.M. Burke, et al., JET Collaboration, Phys. Rev. C 90 (1) (2014) 014909.
- [52] P.F. Kolb, U.W. Heinz, Hydrodynamic description of ultrarelativistic heavy ion collisions, in: R.C. Hwa, X.-N. Wang (Eds.), Quark-Gluon Plasma 3, World Scientific, Singapore, 2004, p. 634, arXiv:nucl-th/0305084.
- [53] J.E. Bernhard, J.S. Moreland, S.A. Bass, Nucl. Phys. A 967 (2017) 293.
- [54] M. Djordjevic, M. Gyulassy, R. Vogt, S. Wicks, Phys. Lett. B 632 (2006) 81.
- [55] A. Bazavov, et al., HotQCD Collaboration, Phys. Rev. D 90 (2014) 094503.
- [56] M. Wilde, ALICE Collaboration, Nucl. Phys. A 904-905 (2013) 573c.
- [57] A. Dainese, Eur. Phys. J. C 33 (2004) 495.
- [58] C. Loizides, J. Kamin, D. d'Enterria, Phys. Rev. C 97 (5) (2018) 054910.
- [59] M. Djordjevic, D. Zigic, M. Djordjevic, J. Auvinen, arXiv:1805.04030 [nucl-th].
- [60] Yu. Maezawa, et al., WHOT-QCD Collaboration, Phys. Rev. D 81 (2010) 091501. [61] A. Nakamura, T. Saito, S. Sakai, Phys. Rev. D 69 (2004) 014506.
- [62] S. Acharya, et al., ALICE Collaboration, Phys. Lett. B 788 (2019) 166.
- [63] ATLAS Collaboration, ATLAS-CONF-2018-007.
- [64] CMS Collaboration, CMS-PAS-HIN-18-004.
- [65] D. Molnar, D. Sun, Nucl. Phys. A 932 (2014) 140;
- D. Molnar, D. Sun, Nucl. Phys. A 910-911 (2013) 486. [66] T. Renk, Phys. Rev. C 85 (2012) 044903.
- [67] M. Djordjevic, B. Blagojevic, L. Zivkovic, Phys. Rev. C 94 (4) (2016) 044908.
- [68] J. Mihee, CMS Collaboration, Nucl. Phys. A 904-905 (2013) 657c.
- [69] J. Xu, J. Liao, M. Gyulassy, Chin. Phys. Lett. 32 (2015) 092501; S. Shi, J. Liao, M. Gyulassy, Chin. Phys. C 42 (10) (2018) 104104.